

Numerical Integration

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Like in root finding, not all integrals have analytic answers. There are two main cases we will look at here. The first is you have a function with a known form and the second is when you just have data points (or function evaluations).

As an example we may want to know the fraction of the Sun's energy radiated in the visible part of the solar spectrum. The amount of energy as a function of wavelength is given by

$$B_\lambda = \frac{2hc^2}{\lambda^5} \left(\frac{1}{e^{hc/\lambda kT} - 1} \right) \quad (1)$$

where λ is the wavelength, h is Planck's constant, k is Boltzmann's constant, c is the speed of light, and T is the temperature. Stating our above question mathematically,

$$f_E = \frac{\int_{400nm}^{700nm} B_\lambda d\lambda}{\int_0^\infty B_\lambda d\lambda} \quad (2)$$

The bottom integral can be done analytically but the top one cannot.

1 Midpoint and Quadrature

The easiest to understand form of numerical integration is using the midpoint rule you may have learned in intro calculus. The idea is simply to evaluate our function at a series of points across the domain. This is also called the rectangle method as we are just summing the area of rectangles. We can write this mathematically as

$$\int_a^b f(x)dx \approx \frac{b-a}{N} \sum_{i=1}^N f(x_i) \quad (3)$$

The more points we evaluate evaluate our function at (larger N) the better estimate we will get.

We can make this method one step better by summing trapezoids instead of rectangles. This method is called trapezoidal rule. There are two main functions that use this method. One is for when you have a function (*quad*) and the other if for when you have data (*trapz*).

```
1
2 import matplotlib.pyplot as plt
3 import numpy as np
4
```

```

5
6 import scipy.optimize as opt
7 import scipy.integrate as integ
8
9 # define our function
10 def B_lmb(wavelength):
11     # Plank function
12     # Returns values in units of W/m^2 nm^-1 sr^-1
13     # wavelength given in nm
14     # Calculates the power power per wavelength
15     h=6.626E-34 # Plank's constant (J/s)
16     c=2.9979E8 # speed of light (m/s)
17     k= 1.380065E-23 # Boltzmann's constant (J/K)
18     T= 5777 # Temperature of the sun (K)
19     wavelength=wavelength*1E-9
20     return (1E-9*2*h*c**2/wavelength**5*
21            (1.0/(np.exp(h*c/(wavelength*k*T))-1)))
22
23 # integrate
24 # the quad function retruns both the result and an estimate of the
25 # error
26 VisE, VisErr=integ.quad(B_lmb, 400, 700)
27
28 # we can use infinity as a limit
29 # Because this function peak sharply at small values I have split
30 # it in two
31 # Try doing it all at once to see the result
32
33 AlleE, AllErr=integ.quad(B_lmb, 0, np.Inf)
34
35 print(VisE/AlleE)

```

Note that here I had the user input the wavelength in nanometers and then convert it to meters. This has no effect when setting limits in quad however when integrating from $0 \leftrightarrow \infty$ quad is more stable if range of interest is between $1 - 10^4$ instead of $10^{-9} - 10^{-5}$.

The sun is well approximated by the Plank function, but that does not take into account the spectral features of the sun. We can try to estimate the same integrals using data of the suns spectrum.

```

1
2 import numpy as np
3 import scipy.integrate as integ
4
5 filename='DataFiles/ASTMG173.csv'
6
7 Data=np.loadtxt(filename, skiprows=2, delimiter=',')
8
9 wvl= Data[:,0] # wavelength in nm
10 Energy = Data[:,1] # Energy at the Earth above the atmosphere (W/(m
11     ^2 nm))
12
13 AlleE = integ.trapz(Energy, x=wvl) # integrate over all wavelengths
14     in file
15
16 wl_max=700 # nm
17 wl_min=400 # nm
18 Vis_index = np.logical_and(wvl>=wl_min, wvl<=wl_max) # Setup
19     logical index for visible range

```

```

19 VisE = integ.trapz(Energy[Vis_index],x=wvl[Vis_index])
20
21
22
23 print(VisE/AlIE)

```

These methods differ in their answers by $\sim 3\%$ (check for yourself) but both agree that $\sim 40\%$ of the sun's light is in the visible range.

2 Monte-Carlo

If you have a function, you can also estimate the integral of that function using a set of random numbers. To do this we do need to know the limits of integration, and the limits of the function on that domain. The general idea here is to randomly throw points within a square domain (in x and y) You can tally how many of those points fall below your curve and how many are above. This gives an estimate for the fraction of that domain that is occupied by your function It is easy to calculate that area of the box, and from that you get the area occupied by your function. Below is an example

```

1
2 import matplotlib.pyplot as plt
3 import numpy as np
4
5 import numpy.random as rnd
6
7 import scipy.optimize as opt
8 import scipy.integrate as integ
9
10 import time # for checking speed
11
12 # define our function
13 def B_lmb(wavelength):
14     # Plank function
15     # Returns values in units of W/m^2 nm^-1 sr^-1
16     # wavelength given in nm
17     # Calculates the power power per wavelength
18     h=6.626E-34 # Plank's constant (J/s)
19     c=2.9979E8 # speed of light (m/s)
20     k= 1.380065E-23 # Boltzmann's constant (J/K)
21     T= 5777 # Temperature of the sun (K)
22     wavelength=wavelength*1E-9
23     return(1E-9*2*h*c**2/wavelength**5*
24            (1.0/(np.exp(h*c/(wavelength*k*T))-1)))
25
26
27
28 # plot up function
29 Plotx=np.linspace(0,5000,100)
30
31 plt.figure()
32 plt.plot(Plotx, B_lmb(Plotx), 'k')
33
34 Ndarts=int(1E4) # number of mote-carlo darts to throw for
35                 # integration
36 xlimits=[400.0,700.0] # The domain we want to integrate over
37 ylimits=[0.0,40.0E3] # the maximum of the function over this range
38                 # is ~26E3

```

```

38 TotalArea = (xlimits[1]-xlimits[0])*(ylimits[1]-ylimits[0])
39
40 # I will use a loop method and vector method and check which is
    faster
41
42 Lt0 = time.time()
43 # Loop method
44 countbelow=0
45 for i in range(Ndarts):
46     # Generate a random x-y point in the range above
47     x= rnd.rand()*(xlimits[1]-xlimits[0])+xlimits[0]
48     y= rnd.rand()*(ylimits[1]-ylimits[0])+ylimits[0]
49
50     if y<B.lmb(x): # if it is below our function
51         countbelow+=1 # count it
52         plt.plot(x,y, 'b.') # plot it blue
53     else:
54         plt.plot(x,y, 'r.') #otherwise plot it red
55         #pass # if you comment out the plot you need pass to have
    something here
56
57 IntArea=float(countbelow)/float(Ndarts)*TotalArea # float function
    keeps integer math from messing this up
58 Lt1 = time.time()
59
60 #Vector method
61 # make arrays of random x-y points
62 Vt0 = time.time()
63 xarr= rnd.rand(Ndarts)*(xlimits[1]-xlimits[0])+xlimits[0]
64 yarr= rnd.rand(Ndarts)*(ylimits[1]-ylimits[0])+ylimits[0]
65
66 BelowIndex=yarr<B.lmb(xarr) #create a logical array for who is
    below the curve
67
68 countbelow2=np.sum(BelowIndex) # Trues count as 1, False is 0
69
70
71
72 IntArea2=float(countbelow2)/float(Ndarts)*TotalArea # float
    function keeps integer math from messing this up
73
74 #make new figure
75 plt.figure()
76 plt.plot(Plotx, B.lmb(Plotx), 'k')
77
78 plt.plot(xarr[BelowIndex],yarr[BelowIndex], 'b.') # plot it blue if
    below
79 plt.plot(xarr[np.logical_not(BelowIndex)],yarr[np.logical_not(
    BelowIndex)], 'r.') #otherwise plot it red
80
81 Vt1=time.time()
82
83 plt.xlabel('Wavelength (nm)')
84 plt.xlabel('Energy ( W/m^2 nm^-1 sr^-1)')
85
86 # integrate with quad for comparison
87 # the quad function retruns both the result and an estimate of the
    error
88 VisE, VisErr=integ.quad(B.lmb, 400, 700)
89
90 # Compare answers
91 print('Quad\tMonte Carlo 1\tMonte Carlo 2')

```

```
92 print('{:0.3E}\t{:0.3E}\t{:0.3E}'.format(VisE, IntArea, IntArea2))
93 print('')
94 print('Speed check (s)')
95 print('Loop\tVector\tDifference')
96 Ldt=Lt1-Lt0
97 Vdt=Vt1-Vt0
98 print('{:0.3f}\t{:0.3f}\t{:0.3f}'.format(Ldt, Vdt, Ldt-Vdt))
```

In the above example I do the Monte-Carlo integration as a loop and as a vector process. In my test I found the vector method was **ninety times faster**. I commented out all the plotting commands to try and be fair to the loop in my test.